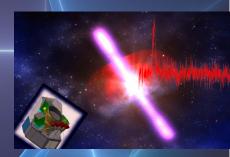
Dorottya Szécsi

Humboldt Fellow University of Cologne



Symposium S4, EWASS, Lyon 24th June 2019



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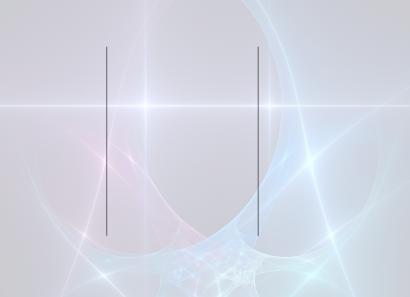
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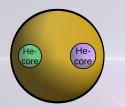
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Common envelope

Chemicallyhomogeneous evolution in a binary

Dynamics in dense clusters

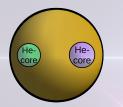
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Common envelope in a binary Chemicallyhomogeneous evolution in a binary

Dynamics in dense clusters

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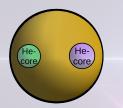


Common envelope in a binary 00

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Common envelope in a binary

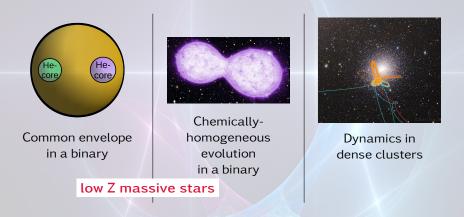


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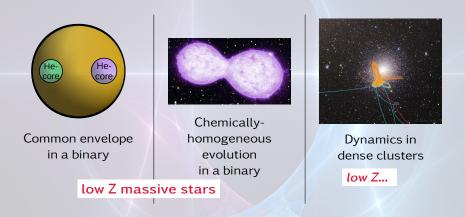


Dynamics in dense clusters

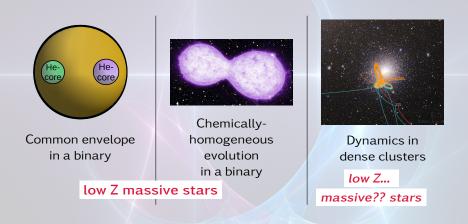
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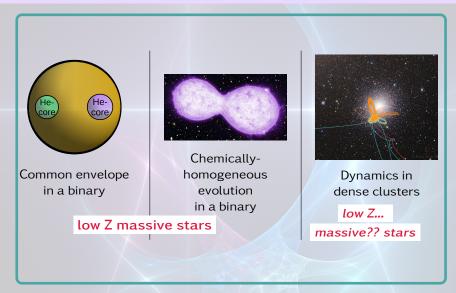
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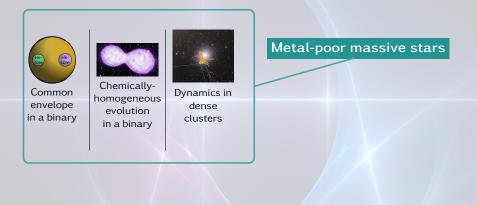
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e.g. Vigna-Gómez..<u>Szécsi</u>+18; <u>Szécsi</u>'17a,b; <u>Szécsi</u>&Wünsch'18; <u>Szécsi</u>'16;

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Chemically homogeneous evolution as a **single** star Chemicallyhomogeneous evolution in a binary

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Chemically homogeneous evolution as a **single** star Chemicallyhomogeneous evolution in a binary

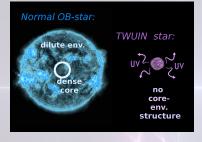
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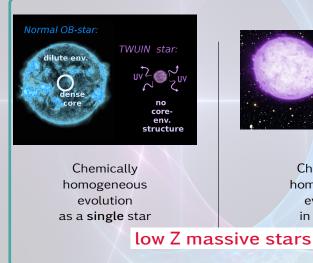




Chemically homogeneous evolution as a **single** star Chemicallyhomogeneous evolution in a binary

low Z massive stars

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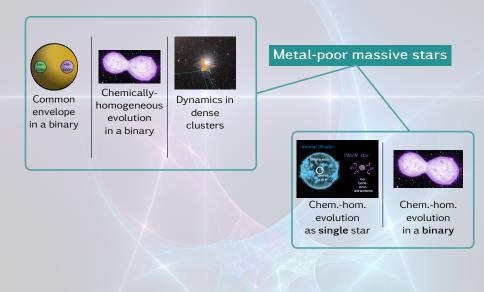


e.g. Yoon&Langer'05; Woosley&Heger'06; Yoon+06; <u>Szécsi</u>+15; <u>Szécsi</u>'16; Marchant+16; <u>Szécsi</u>'17a,b

Chemicallyhomogeneous evolution in a binary

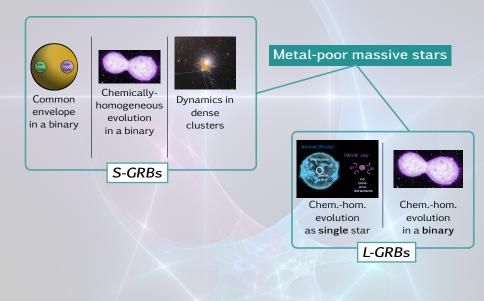
GRB progenitors

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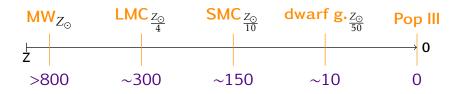
However...

Are they observed?

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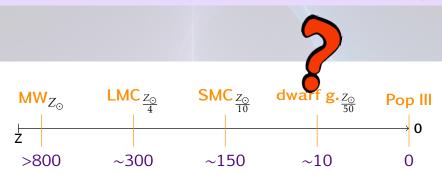


spectroscopy (i.e. direct evidence)

e.g. Castro+14,+18, Ramírez-Agudelo+17, Kubátová&Szécsi+18

Are they observed?

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spectroscopy (i.e. direct evidence)

GRB-progenitors theories...

e.g. Castro+14,+18, Ramírez-Agudelo+17, Kubátová&Szécsi+18

Indirect evidence!

Dwarf galaxies

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Legrand+07, Aloisi+09, Annibali+13, Kehrig+13, Lebouteiller+13

Dwarf galaxies

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I Zwicky 18's ionization

- observed: $Q_{He II} = 10^{50} \gamma/s$ Kehrig+2015
- explained by Chem. Hom. single stars
 Szécsi+15
- explained by Pop-III stars e.g. Heap+16
- explained by X-ray binaries Schaerer+19



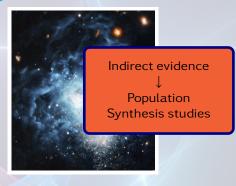
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a synergy between GRBs & massive stars





- e.g.: implement rotating stellar models into population synthesis codes
- e.g.: implement collapsar conditions into stellar evolution code MESA



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- train new PhDs in joint projects
- use existing GRB data \rightarrow prepare for THESEUS
- hold conferences to build connections between GRB & massive star communities