Globular Cluster Abundance Anomalies and the Massive Binary Polluter Scenario

Dorottya Szécsi Nicolas Conzalez-Jimenez Norbert Langer Argelander-Institut für Astronomie

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A grid of low metallicity single stars



Szécsi et al. 2014 in prep.

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Abundance anomalies observed in Galactic Clusters (GCs)

















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 - massive binaries: non-conservative mass transfer (De Mink+ 2009)

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- *De Mink*+ 2009: **20** M_{\odot} + **15** M_{\odot} + **12** days (~0.025 Z_{\odot})









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$$RL_1 = A \frac{0.49q^{2/3}}{0.6q^{2/3} + \ln(1+q^{1/3})}$$



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⇒ when $R_1 \ge RL_1$: check size and composition of the primary envelope

Composition and size of primary envelope



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- give constraints on the massive binary polluter scenaro even without detailed binary simulations











Compared to observations:

O – Na anticorrelation









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Mg - Al anticorrelation




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 - extended for higher masses (up to \sim 575 M_{\odot})

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Thank you for your attention!