

Part I: Consequences of Mass Loss on the Final Fates of Massive Stars

Part II: The BEC Interface

Dorottya Szécsi



Group Meeting, Bonn

12. June 2014

Stellar evolution + mass loss recipes

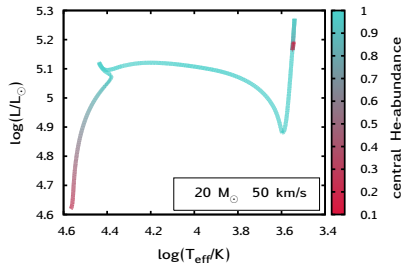
Stellar evolutionary tracks

- hydrodynamic simulation of an isolated, rotating gas sphere (= **star**)
- nuclear burning, 1D

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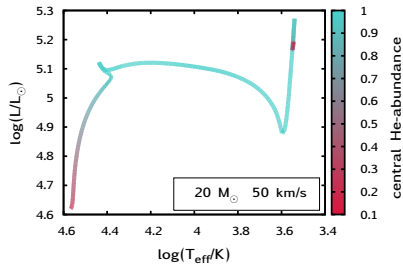


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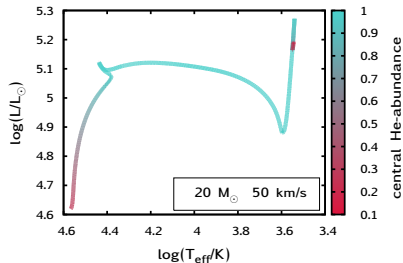
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- model atmospheres with different L_* , M_* , T_{eff} , v_∞/v_{esc} (Vink et al. 2000)
- OR spectral analyses



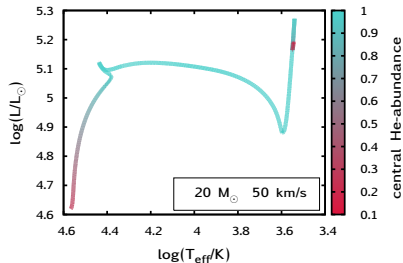
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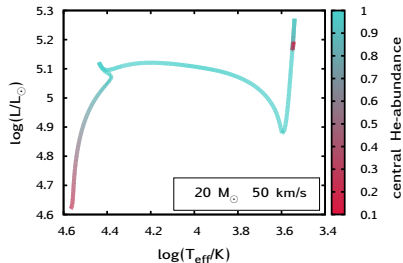
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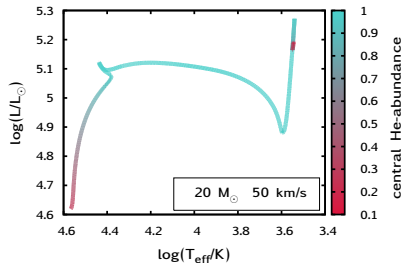
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e.g. Hamann et al. 1995 (for $\log(L/L) > 4.5$):

$$\log \dot{M} = -11.95 + 1.5 \log \frac{L_*}{L} + 2.85 X_s + 0.86 \log Z$$

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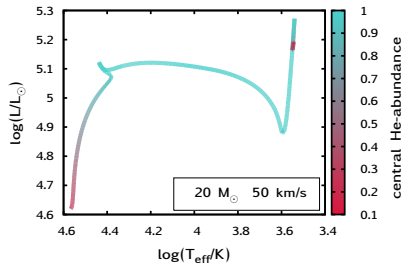
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- \dot{M} is calculated in every step and the corresponding M is removed

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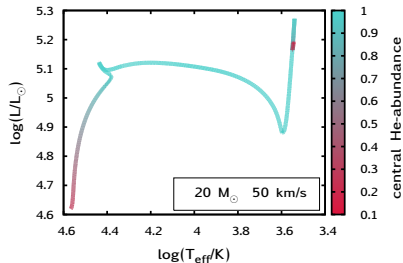
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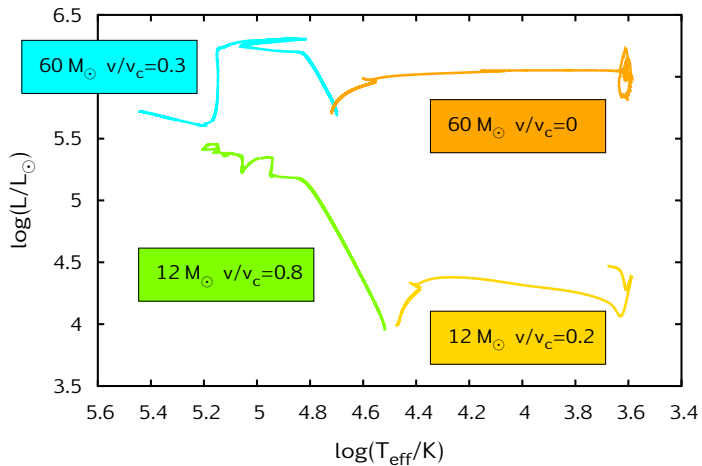
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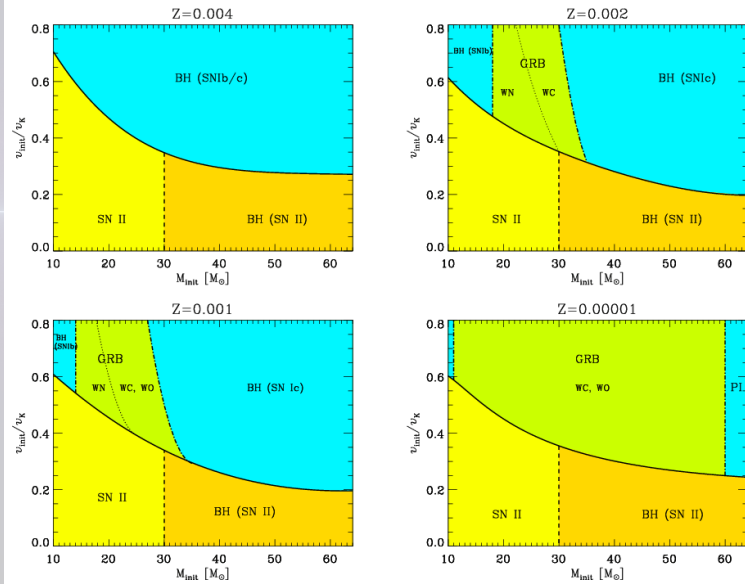
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 \rightarrow mass loss rate has a feedback on the evolution!

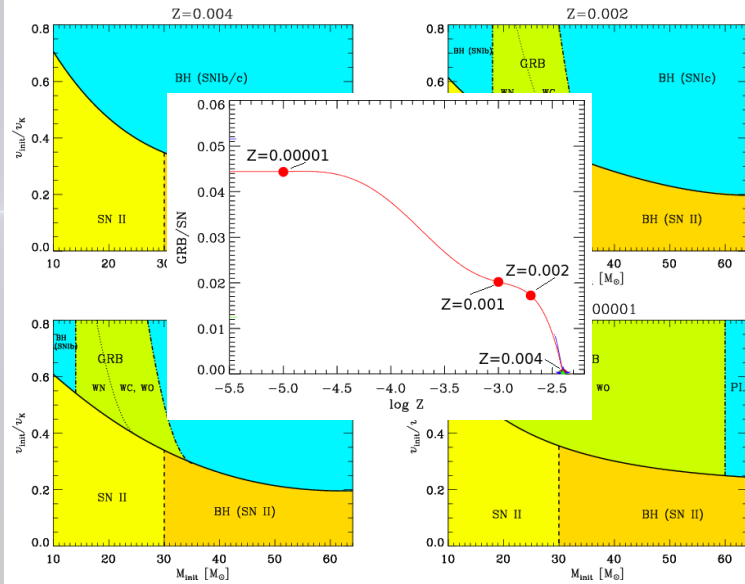
Yoon et al. 2006: low Z tracks on the HRD



Yoon'06: IGRB and SN progenitors at different Z



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The infamous Wolf–Rayet mass loss

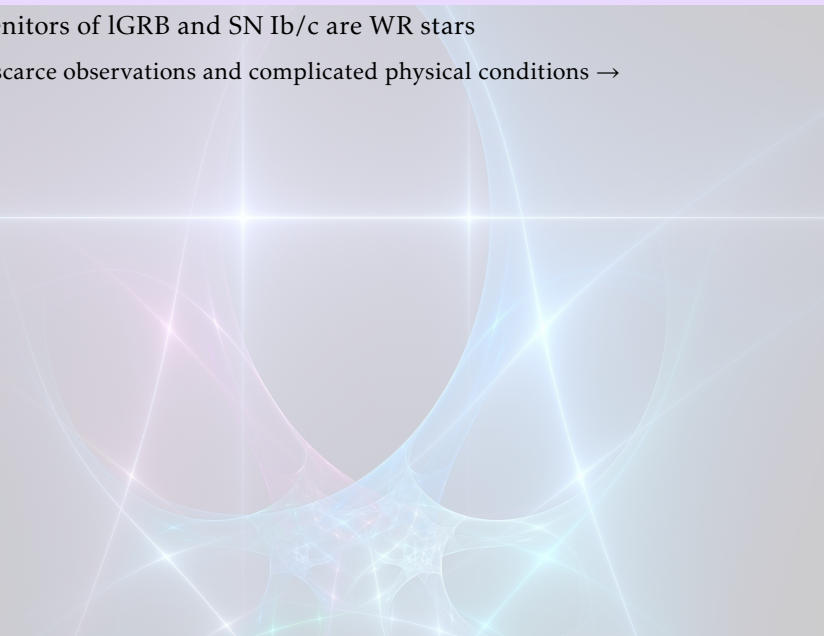
Progenitors of IGRB and SN Ib/c are WR stars



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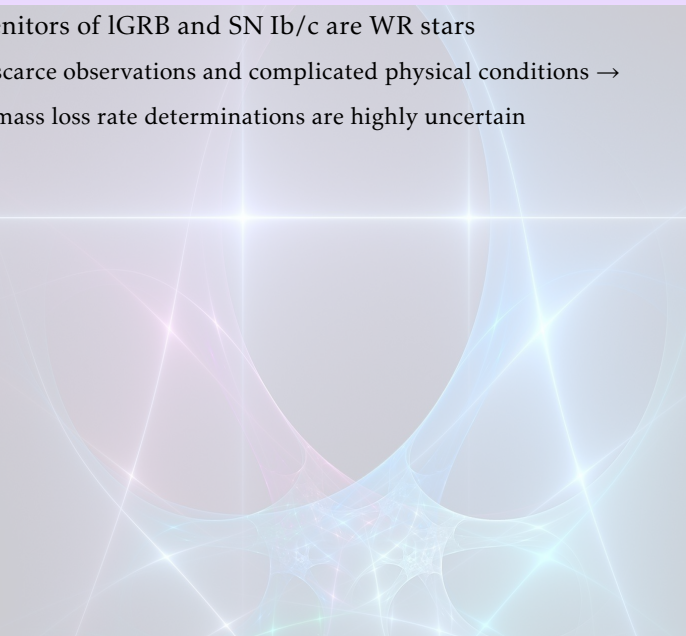
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...and on the final fate too!

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- enhanced mass loss due to CNO in the surface: $\dot{M} = f \cdot \dot{M}_{H95}$,
 $f \sim 19 \cdot Z_{CNO}^{surf}$

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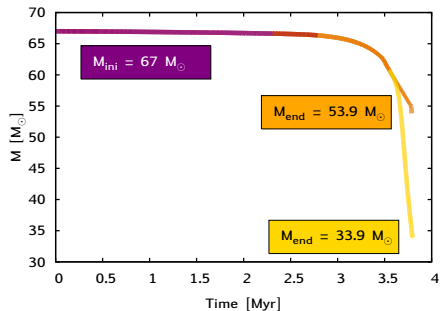
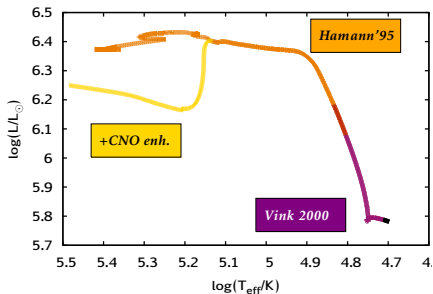
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 - How much does it effect the final fate predictions?

With and without CNO enhanced mass loss

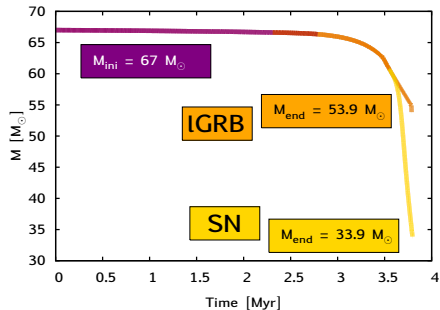
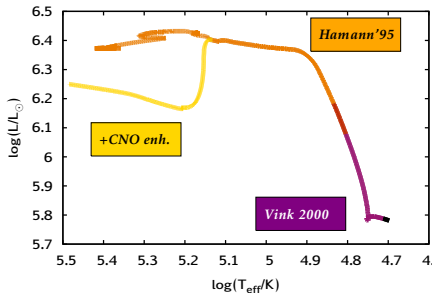
67 M_{\odot} $v/v_c=0.3$ tracks from Szécsi et al. 2014



- IGRB: fast rotating WR star (collapsar model)
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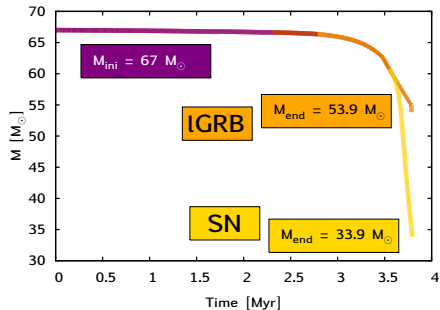
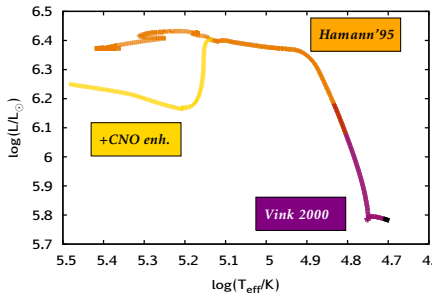
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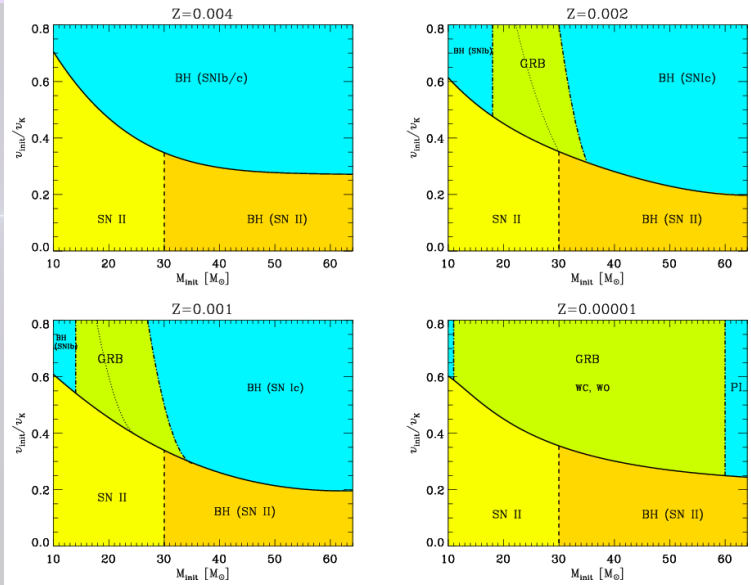
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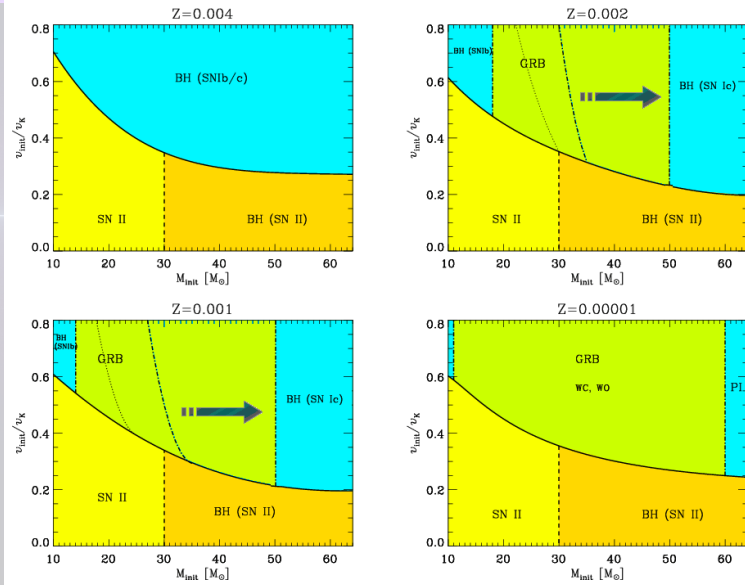


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- If CNO enh. massloss is unreasonable: more IGRBs and less SNe \rightarrow

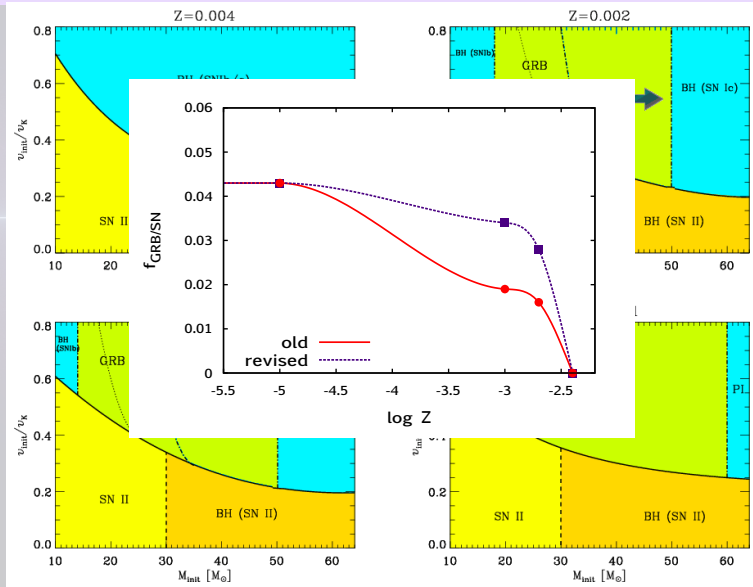
I GRB rate of Yoon'06 – REVISED



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Which WR mass loss prescription is more valid?

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- Waiting for comments and questions!



Thank you for
your attention!

WR wind mass loss rates

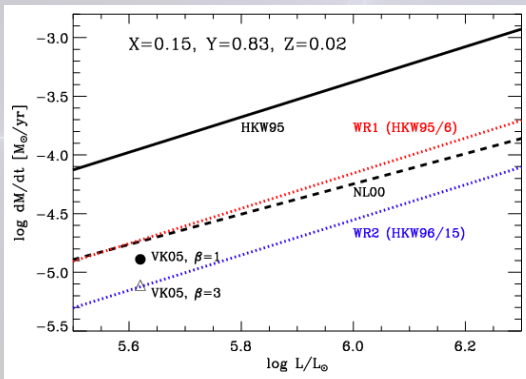


Fig. 1. of Yoon & Langer 2005

Wolf-Rayet wind mass loss rates as a function of the stellar luminosity for a given surface composition.

Hamann et al. 1995: HKW95 (solid)

Nugis & Lamers 2000: NL00 (dashed)

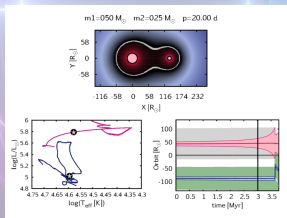
HKW/6

HKW/15

Vink & de Koter 2005: VK05
(mass loss rate for WN stars)

The BEC interface

Dorottya Szécsi
Argelander-Institut für Astronomie



June 12, 2014

The BEC interface: `beci`

BEC = Binary Evolutionary Code

- single and binary stellar systems
- with or without rotation
- developed since the 1970s in Fortran
- very powerful!
- ...but difficult to use

`beci` = the interface for BEC

- automatized + documented
- easy to learn and use
- transferable, extensible
- and more...

Basic commands 1.

Run a new single stellar evolutionary model

```
WORKDIR/BEC$ ./beci -single 20 0.8 50 1mc -run
```

$M_{ini}=20 M_{\odot}$ $Z=0.8*Z_{LMC}$ $v_{ini}=50$ km/s

Check its status

```
./beci -single 20 0.8 50 1mc -v1
```

Create a HRD

```
./beci -single 20 0.8 50 1mc -visualize -HRD
```

Create a Kippenhahn diagram

```
./beci -single 20 0.8 50 1mc -visualize -kippenhahn
```

Basic commands 2.

Get help any time

```
./beci -help  
./beci -help -run
```

Modify the m.dat from the command line

```
./beci -single 20 0.8 50 lmc -mdat IOUT=20 DTMIN_1=1.0d0
```

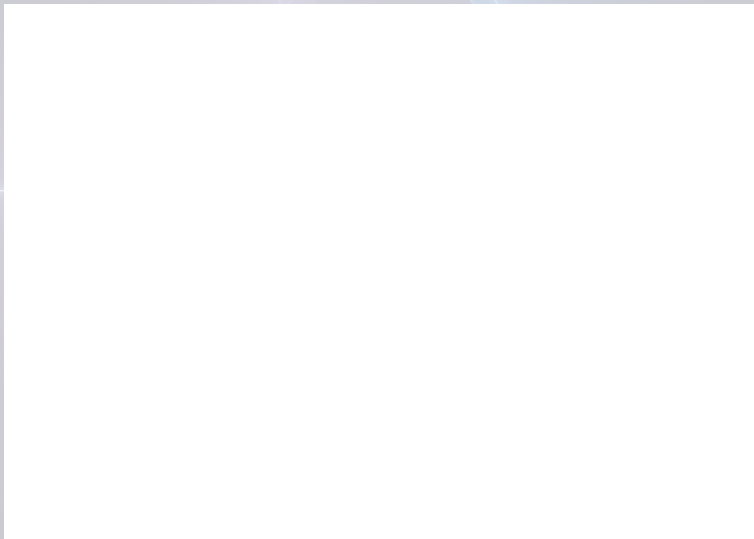
Continue a stopped calculation

```
./beci -single 20 0.8 50 lmc -continue
```

Create an animation of the composition change

```
./beci -single 20 0.8 50 lmc -visualize -composition
```

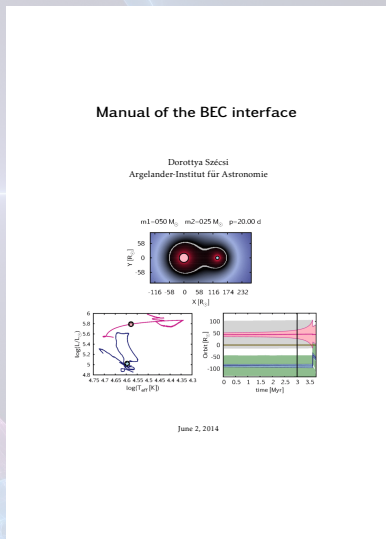

Basic commands: -composition



More options for single stars in the Manual

- -normal
- -extended
- -savetozams
- -comeclean
- -obsHRD
- -recipes
- -structure
- -elements
- -YcYs
- -yield
- -angmom
- -PISN

Also see the -help
command!



So what about binaries?

Run a new binary stellar evolutionary model

```
./beci -binary 050 025 1.0 20.00 smc -run
```

$$M_{ini}^P = 50 M_{\odot} \quad M_{ini}^S = 25 M_{\odot} \quad Z = 1.0 * Z_{SMC} \quad p_{ini} = 20 \text{ days}$$

Check their status

```
./beci -binary 050 025 1.0 20.00 smc -v1 -v2 -v3
```

Visualize them on the HRD and Orbit diagram

```
./beci -binary 050 025 1.0 20.00 smc -visualize
```

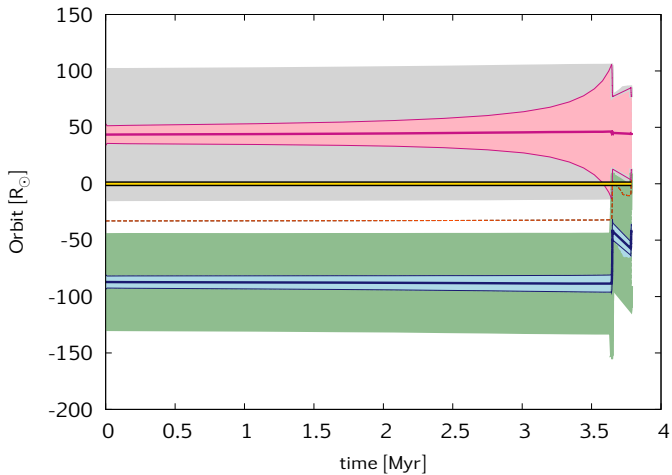
More options for binary stars in the Manual:

- -mdat -continue -normal -kippenhahn -composition
-recipes -structure -help

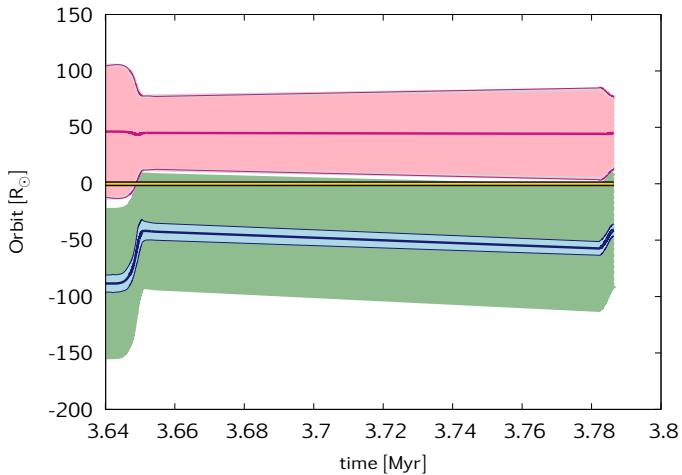
...and more

```
./beci -binary 050 025 1.0 20.00 smc -visualize -anim
```

The 'Orbit' diagram



The 'Orbit' diagram - after RLOF



Handling large grids of stars

Grids created by BEC:

- Yoon et al. 2006, 2012 (low Z, Pop.III)
- Brott et al. 2011 (MW, LMC, SMC)
- Köhler et al. 2014 (LMC)
- Szécsi et al. 2014 (IZw18)

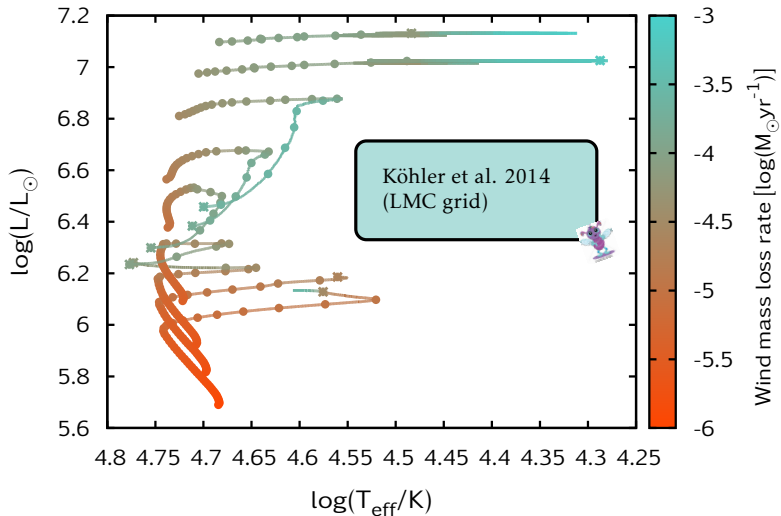
Common HRD of all stars in the grid

```
./beci -grid NameOfGrid -gridHRD
```

More grid commands are available, see the Manual or call

```
./beci -help -grid
```

Handling large grids of stars: -gridHRD



Write your own script!

```
./beci -single 20 0.8 50 lmc -myscript -coolscript Dori
```

```
##coolscript.sh
echo "Hi dear friend $!"
echo "You are working with the star
M=$m Msun - z=$z x Z_$g - v=$v km/s in $path_BEC/bin"
endMS=`cat ./m$m-z$z-v$v.$g/$m-$z.plot1 | sed 's/D/E/g' | awk '
BEGIN{max=0}{if($3>max){max=$3;tmax=$1}} END{print tmax}`
endMSMyr=`echo $endMS | awk '{printf "%1.3f", $1/1000000 }`
echo "Main sequence lifetime: $endMSMyr Myr"
echo "Bye! :)"
```

Result

```
Hi dear friend Dori!
You are working with the star
M=20 Msun - z=0.8 x Z_lmc - v=50 km/s in $WORKDIR/BEC/bin
Main sequence lifetime: 7.886 Myr
Bye! :)
```

Thank you for your attention!

Program files are available here:

```
/vol/cstorage/raid18/dorottya/BECinterface
```

Copy the files in your \$WORKDIR and call

```
./beci -setup
```



Please try it and
find bugs!