



# Skeleton Hamiltonian dynamics

*Mathematics of Gravitation II  
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- A Wilson-like approach





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- Standard Hamiltonian description of  $N$  black holes





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- Application to circular motion





# A WILSON-LIKE APPROACH



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**Objective:** finding an analytical counterpart of Wilson



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  - in the test particle limit
- Can be derived from a Hamiltonian  
    → investigations performed in the ADM framework



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- gives the correct Brill-Lindquist solution in the limit  $p_{Ai} = 0$
- supported by the “effacement” principle



# STANDARD ADM DESCRIPTION



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- $(g_{0i}g_{0j}\gamma^{ij} - g_{00})^{\frac{1}{2}} = N$  = lapse  
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## Introduction of the canonical variables:

- $(x_A^i, p_{Ai})$  = matter conjugate variables
- $(\gamma_{ij}, \pi^{ij})$  = field conjugate variables





# Hamiltonian for $N$ point-masses

$$H(x_A^i, p_{Ai}, \gamma_{ij}, \pi^{ij}) = \frac{c^4}{16\pi G} \left[ \int d^3x (N\mathcal{H} + N^i \mathcal{J}_i) \right. \\ \left. + \int d^3x \partial_i (\partial_j \gamma_{ij} - \partial_i \gamma_{jj}) \right]$$

with  $\mathcal{H} = -\sqrt{\gamma}R + \frac{1}{\sqrt{\gamma}} \left( \pi_j^i \pi_i^j - \frac{1}{2} (\pi_i^i)^2 \right)$

$$+ \frac{16\pi G}{c^2} \sum_{A=1}^N \left( m_A^2 + \frac{p_{Ai} p_{Ai}}{c^2} \right)^{\frac{1}{2}} \delta_A$$

[super-Hamiltonian]

and  $\mathcal{J}_i = -2\partial_j \pi_i^j + \pi^{kl} \partial_i \gamma_{kl} - \frac{16\pi G}{c^3} \sum_{A=1}^N p_{Ai} \delta_A$

[supermomentum]





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- decomposition of the field variables
  - decomposition of  $\gamma_{ij}$  for a conformal factor  $\Psi^4$ :

$$\gamma_{ij} = \Psi^4 \delta_{ij} + h_{ij}^{\text{TT}} \text{ with } \partial_j h_{ij}^{\text{TT}} = 0$$





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- field momentum decomposition

$$\pi^{ij} = \text{STF}(2\partial_i \pi^j) + \pi_{\text{TT}}^{ij} \text{ with } \partial_j \pi_{\text{TT}}^{ij} = 0$$



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- Final Hamiltonian

$$\begin{aligned} H(x_A^i, p_{Ai}, h_{ij}^{\text{TT}}, \pi_{\text{TT}}^{ij}) &= \frac{c^4}{16\pi G} \int d^3x \, \partial_i (\partial_j \gamma_{ij} - \partial_i \gamma_{jj}) \\ &= -\frac{c^4}{2\pi G} \int d^3x \, \Delta \Psi \end{aligned}$$





# EFFECTIVE SKELETON HAMILTONIAN



# 1<sup>st</sup> ingredient of the model

$\gamma_{ij} \propto \delta_{ij}$ :

← skip the complications due to the  $h_{ij}^{\text{TT}} \neq 0$  dynamics  
lose of accuracy if  $v/c$  large *but* good in the static limit





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- $\pi^{ij} = \gamma^{ik} \pi_k^j = \Psi^{-4} \pi_j^i$   
 $\Rightarrow \pi_i^i = 0$  (maximal slicing)  
 $\pi_j^i$  STF (symmetric trace-free)



# Constraints equations for $h_{ij}^{\text{TT}} = 0$

$$\Delta\Psi = -\frac{1}{8\Psi^7}\pi_j^i\pi_i^j - \frac{2\pi G}{c^2} \sum_{A=1}^N m_A \delta_A \Psi^{-1} \left[ 1 + \frac{p_{Ai}p_{Ai}}{m_A^2 c^2 \Psi^4} \right]^{\frac{1}{2}}$$

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## 2<sup>nd</sup> ingredient: specific form of $\pi_j^i$

$$(\pi_j^i)^{\text{TT}} = \phi(\pi_j^i)^{\parallel} + 2\tilde{\pi}_{(5)}^{ij} + \mathcal{O}\left(\frac{1}{c^7}\right) = \mathcal{O}\left(\frac{1}{c^5}\right) \rightarrow \text{contributes to the 3PN dynamics}$$

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regular at  $r_A = |\mathbf{x} - \mathbf{x}_A| = 0$  for appropriate dimensions  
agreement with Bowen & York





# Flesh and skeleton

$$\begin{aligned}\Psi^{-7} \pi_j^i \pi_i^j &= \Psi^{-7} \pi_j^i \text{STF}(2\partial_i V_j) \\ &= \Psi^{-7} (-2V_i \partial_j \pi_j^i + 2\partial_j(V_i \pi_j^i)) = \mathcal{O}\left(\frac{1}{c^6}\right)\end{aligned}$$





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skeleton                              ↑  
    flesh





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Properties of the flesh terms





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- vanishes when  $p_{Ai} = 0$
- vanishes in the test body limit
- contains **poles** in dimensional regularisation  
  ← compensated by poles in  $h_{ij}^{\text{TT}}$  in Einstein theory



# 3<sup>rd</sup> ingredient: $2\partial_j(V_i\pi^i_j) \rightarrow 0$

**elimination of the flesh term**

$$\Psi^{-7}\pi^i_j\pi^j_i \rightarrow \Psi^{-7}(-2V_j\partial_i\pi^i_j) = \frac{16\pi G}{c^3} \sum_{A=1}^N \Psi_{x=x_A}^{-7} p_{Aj} V_j \delta_A$$

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structure of  $\Psi$

$$\left\{ \begin{array}{l} \Delta\Psi = \sum_A f_A(x)\delta_A \\ \Psi = 1 + \mathcal{O}\left(\frac{1}{r}\right) \text{ at spatial infinity} \end{array} \right. \Rightarrow \Psi = 1 + \sum_{A=1}^N \frac{G\alpha_A}{2r_A c^2}$$



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**skeleton super-Hamiltonian constraint**  
→ obtained by inserting  $\phi$  into  $\mathcal{H}$





# Hamiltonian constraint

$$\begin{aligned} -\frac{c^2}{2\pi G} \Delta \Psi &= \sum_{A=1}^N \alpha_A \delta_A \\ &= \sum_{A=1}^N \left\{ m_A \left( 1 + \sum_{B \neq A} \frac{G\alpha_B}{2r_{AB}c^2} \right)^{-1} \left[ 1 + \left( 1 + \sum_{C \neq A} \frac{G\alpha_C}{2r_{AC}c^2} \right)^{-4} \frac{p_{Ai}p_{Ai}}{m_A^2 c^2} \right] \right. \\ &\quad \left. + \left( 1 + \sum_{B \neq A} \frac{G\alpha_B}{r_{AB}c^2} \right)^{-7} \frac{p_{Ai}V_{Ai}}{c} \right\} \delta_A \end{aligned}$$





# Equation for the lapse

- { evolution of  $\pi^{ij}$  in a (3+1) splitting:  $\partial_t \pi^{ij} = F^{ij}[\gamma_{kl}, N_k, x_A^k, p_{Ak}]$
- { gauge condition:  $\pi^{kk} = 0$
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⇓

$$\begin{aligned} N\Delta\Psi + \Delta [\Psi N] &= \frac{3}{4}\Psi^{-7}N\pi_j^i\pi_i^j \\ &+ \frac{4\pi G}{c^4}\Psi^{-5}N\sum_{A=1}^N \delta_A \frac{p_{Ai}p_{Ai}}{m_A} \left(1 + \frac{p_{Ai}p_{Ai}}{m_A^2 c^2 \Psi^4}\right)^{1/2} \end{aligned}$$





# Equation for the shift function

Evolution equation for  $\gamma_{ij}$

$$\frac{1}{c} \partial_t \gamma_{ij} = \frac{N}{\sqrt{\gamma}} (2\pi_{ij} - \pi_k^k \gamma_{ij}) + 2D_{(i} N_{j)}$$

with  $2D_{(i} N_{j)} = \gamma_{ki} \partial_j N^k + \gamma_{kj} \partial_i N^k + N^k \partial_k \gamma_{ij}$





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$$\gamma_{ij} - \frac{1}{3} \gamma_{kk} \delta_{ij} = 0 \quad \implies \quad \partial_t \left( \gamma_{ij} - \frac{1}{3} \gamma_{kk} \delta_{ij} \right) = 0$$





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in particular:  $\Delta N^i + \frac{1}{3} \partial_{ij} N^j = -2 \partial_j (\Psi^{-6} N \pi_j^i)$



# Lapse function

**Auxiliary quantities** natural to pose:  $N = \chi/\Psi$  with  $\lim_{|x| \rightarrow +\infty} N = 1$





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$$\chi = 1 - \sum_{A=1}^N \frac{G\beta_A}{2r_A c^2}$$





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$$\begin{aligned} \Psi^{-1} \chi \Delta \Psi + \Delta \chi &= \frac{3}{4} \Psi^{-8} \chi \pi_j^i \pi_i^j \\ &+ \frac{4\pi G}{c^4} \Psi^{-6} \chi \sum_{A=1}^N \delta_A \frac{p_{Ai} p_{Ai}}{m_A} \left( 1 + \frac{p_{Ai} p_{Ai}}{m_A^2 c^2 \Psi^4} \right)^{1/2} \end{aligned}$$





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$$\Psi^{-1} \chi \Delta \Psi + \Delta \chi = \frac{3}{4} \Psi^{-8} \chi \pi_j^i \pi_i^j$$

replaced by

$$\frac{12\pi G}{c^3} \sum_{A=1}^N \left( \frac{\chi}{\Psi^8} \right)_{x=x_A} p_{Ai} V_j \delta_A$$

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# Skeleton Hamiltonian for 2 black holes

equations satisfied by  $\Psi_1 \equiv \Psi_{x=x_1}$  and  $\Psi_2$   
deduced from the linear independence of the  $\delta_A$ 's

$$\Psi_1 = 1 + \frac{Gm_2}{2r_{12}c^2\Psi_2} \left( 1 + \frac{p_2^2}{m_2^2c^2\Psi_2^4} \right)^{\frac{1}{2}} + \frac{Gp_{2i}V_{2i}}{2r_{12}c^3\Psi_2^7}$$

$$\Psi_2 = 1 + \frac{Gm_1}{2r_{12}c^2\Psi_1} \left( 1 + \frac{p_1^2}{m_1^2c^2\Psi_1^4} \right)^{\frac{1}{2}} + \frac{Gp_{1i}V_{1i}}{2r_{12}c^3\Psi_1^7}$$

where  $\Psi_1 = 1 + \frac{G\alpha_2}{2r_{12}c^2}$  and  $r_{12} = |\mathbf{x}_1 - \mathbf{x}_2|$

$$H = -\frac{c^4}{2\pi G} \int d^3x \Delta \Psi = c^2(\alpha_1 + \alpha_2)$$



# Skeleton field for 2 black holes I

3-metric

$$\gamma_{ij} = \Psi^4 \delta_{ij} \quad \text{with} \quad \Psi = 1 + \frac{G}{2c^2} \left( \frac{\alpha_1}{r_1} + \frac{\alpha_2}{r_2} \right)$$



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**Lapse function**

obtained by projection of the lapse equation on  $\delta_A$

$$\begin{aligned} \chi_B &= 1 - \frac{Gm_A}{r_{12}c^2} \Psi_A^{-4} \chi_A \left\{ \frac{7p_{Ai}(V_{i\mathbf{x}=\mathbf{x}_A})}{m_A c} \right. \\ &\quad \left. + \left[ 1 + \frac{p_{Ai}p_{Ai}}{m_A^2 c^2 \Psi_A^4} \right]^{-1/2} \left[ 3\Psi_A^2 \frac{p_{Ai}p_{Ai}}{m_A^2 c^2} + \Psi_A^6 \right] \right\} \end{aligned}$$

$$\text{where } \chi_1 = 1 - \frac{G\beta_2}{2r_{12}c^2}$$

$$\Rightarrow \text{lapse given by } N = \frac{\chi}{\Psi} \quad \text{with} \quad \chi = 1 + \frac{G}{2c^2} \left( \frac{\beta_1}{r_1} + \frac{\beta_2}{r_2} \right)$$



# Skeleton field for 2 black holes (II)

Shift function



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- 2<sup>nd</sup> skeleton approximation: 2PN flesh term neglected in  $N^i$

$$\partial_j \left( \Psi^{-6} N \pi_j^i \right) \rightarrow \Psi^{-6} N \partial_j \pi_j^i = -\frac{8\pi G}{c^3} \Psi^{-6} N \sum_{A=1}^N p_{Ai} \delta_A$$

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- integration of the shift equation  $\sim$  integration of  $\mathcal{J}_i$

$$N^i = \frac{G}{c^3} \sum_{A=1}^N \chi_A \Psi_A^{-7} \left( \frac{1}{2} p_{Aj} \partial_{ij} r_A - 4 p_{Ai} \frac{1}{r_A} \right)$$

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## Komar mass $M$ :

$$N = 1 - \frac{GM}{c^2|x|} + \mathcal{O}\left(\frac{1}{|x|}\right) \Rightarrow M = \frac{1}{2} \sum_{A=1}^2 (\alpha_A + \beta_A)$$





# Equations of motion

Obtained by means of an implicit differentiation

$$\dot{x}_1^i = \frac{c^2}{(1 - \zeta_1 \zeta_2)} [(1 - \zeta_2) \theta_{11}^i + (1 - \zeta_1) \theta_{21}^i]$$

$$\dot{p}_{1i} = -\frac{c^2}{(1 - \zeta_1 \zeta_2)} [(1 - \zeta_2) \eta_1^i - (1 - \zeta_1) \eta_2^i] - \frac{2c^4}{G} (\Psi_1 + \Psi_2 - 2) n_{12}^i$$

with  $\tilde{\eta}_{A1}^i \equiv -\frac{G m_A n_{12}^i}{2c^2 r_{12}^2 \Psi_A} \left( 1 + \frac{p_A^2}{m_A^2 c^2 \Psi_A^4} \right)^{\frac{1}{2}} + \frac{G \partial_{1i} (p_{Aj} V_{Aj} / r_{12})}{2c^3 \Psi_A^7}$

$$\tilde{\eta}_{A2}^i \equiv -\tilde{\eta}_{A1}^i$$

$$\zeta_A \equiv \frac{\chi_B - 1}{\chi_A}$$

$$\tilde{\theta}_{A1}^i \equiv \frac{G p_{1i}}{2m_A c^4 r_{12} \Psi_A^5} \left( 1 + \frac{p_A^2}{m_A^2 c^2 \Psi_A^4} \right)^{-\frac{1}{2}} + \frac{G \partial_{p_{1i}} (p_{Aj} V_{Aj})}{2c^3 r_{12} \Psi_A^7}$$





# CIRCULAR MOTION





# Circular orbits

**PN calculation of  $\hat{H} = \{H - (m_1 + m_2)c^2\}/\mu$ , [ $\mu$  = reduced mass]**

$$\hat{H} = \sum_{A=1}^2 \sum_{i=1}^{n+1} \frac{\alpha_A^{(i)}}{\mu} \varepsilon^i + \mathcal{O}(\varepsilon^{n+2})$$





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- $n_{12}^i p_{1i} = n_{12}^i p_{2i} = 0$ ,
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**theorem: ADM mass = Komar mass**



# Equality of Komar and ADM mass (I)

- dependence of  $\alpha_A$  for circular motion

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- implication of circularity

$$\frac{\partial H(r_{12}, j)}{\partial r_{12}} = 0 \iff \sum_{A=1}^2 \frac{\partial \alpha_A(r_{12}, j)}{\partial r_{12}} = 0$$

$$\iff \sum_{A=1}^2 \left( -\frac{\beta_A}{\chi_A} \frac{\partial \Psi_A}{\partial r_{12}} + \frac{1}{2r_{12}} \left( \alpha_A - \beta_A \frac{\Psi_A}{\chi_A} \right) \right) = 0$$



# Equality of Komar and ADM mass (II)

- expression of  $\partial\Psi_A/\partial r_{12}$

$$\begin{aligned}\Psi_A(r, j) &= 1 + \frac{G\alpha_B}{2r_{12}c^2} \\ \Rightarrow \frac{\partial\Psi_A}{\partial r_{12}} &= -\frac{G}{4r_{12}c^2} \left( \alpha_A + \beta_A \frac{\Psi_B}{\chi_B} \right) + \frac{\beta_B}{\chi_B} \frac{\partial\Psi_B}{\partial r_{12}}\end{aligned}$$

↔ 2 equations for 2 unknowns

solution  $\frac{\partial\Psi_A}{\partial r_{12}} = -\frac{G}{4r_{12}c^2}(\alpha_B + \beta_B), B \neq A$

- insertion in the circularity condition  $\partial H(r_{12}, j)/\partial r_{12} = 0$

result  $\sum_{A=1}^2 (\alpha_A - \beta_A) = 0 \iff M_{\text{ADM}} = M_{\text{Komar}}$

# Application: last stable circular orbit

